

Propagation Forecasting – A Peek Behind the Curtain



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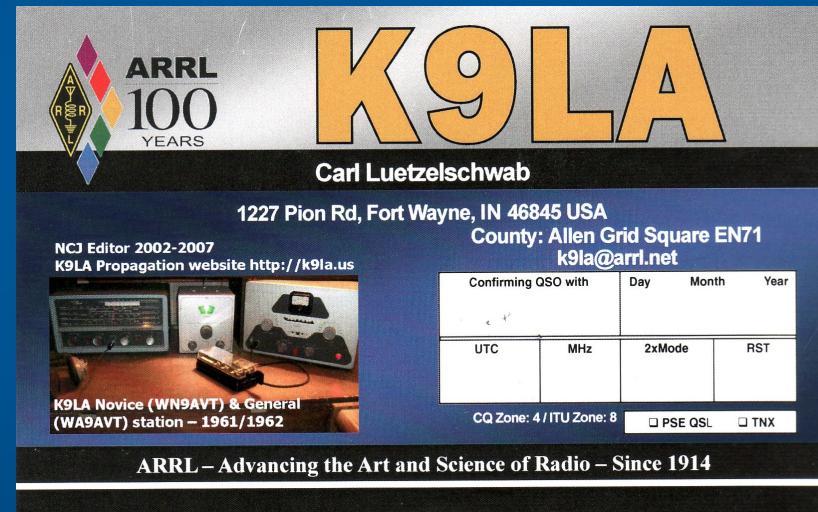


Contrary to popular belief, propagation forecasting is not solely an art – there is science involved. Honest!



Who Is K9LA?

- Licensed in October 1961 as WN9AVT, K9LA in 1977
- Enjoy propagation, DXing, contesting, antennas and vintage equipment
- Interested in propagation since my college days
 - MSEE project about group delay in the ionosphere
- Began doing predictions using the manual method (before PCs)
 - Used worldwide MUF maps, great circle path maps, control points
 - Great way to acquire a fundamental understanding of the process
- Visit <http://k9la.us> for solar and propagation articles



Propagation Predictions

- Propagation predictions (alternatively, propagation forecasts) nowadays refer to using VOACAP or W6ELProp or your favorite software to determine the times and frequencies that will allow you to work a specific target
- This usually gives you a bunch of data
- For an individual trying to increase DXCC or WAZ totals, this is adequate
- But for a contest effort or a DXpedition team, I believe there's more work to do in terms of 'propagation'
 - I call this 'propagation planning'



Common Ground and Agenda

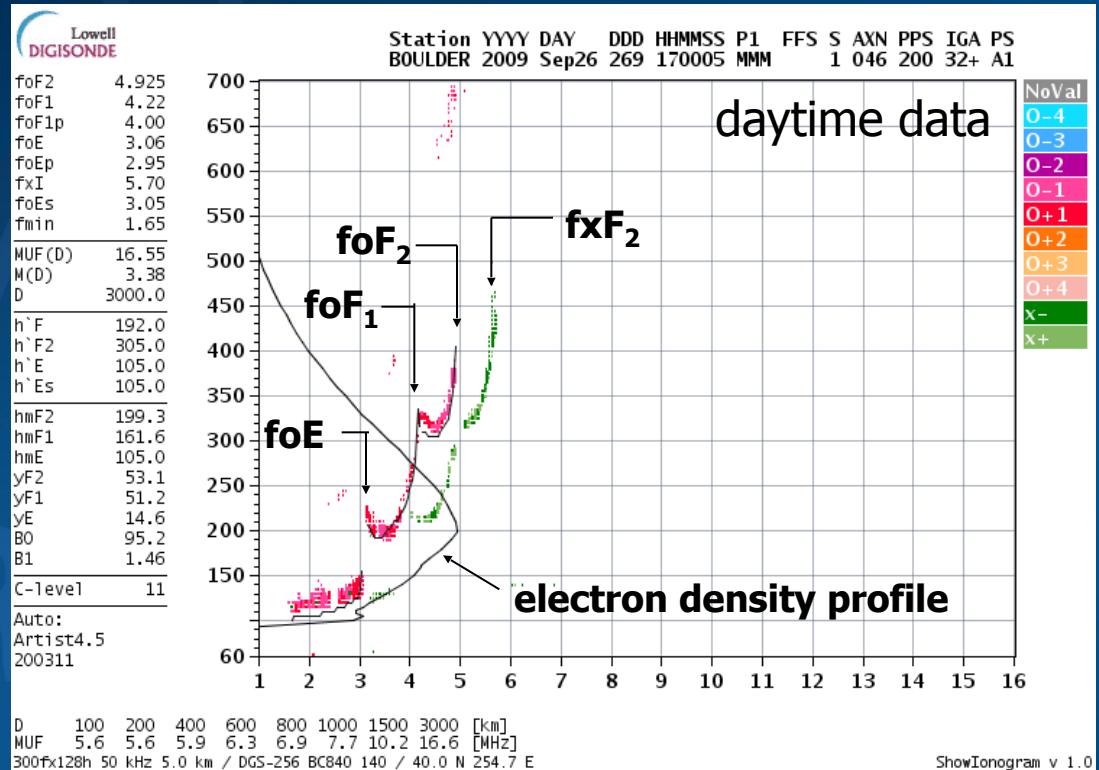
- Regardless of what you do with the predictions, there is common ground – the common ground is the fact that the model of the ionosphere is a monthly median model
- Our agenda will thus be:
 - Quick history and development of the model of the ionosphere
 - How to interpret the results (MUF and signal strength)
 - Predictions for an individual
 - Propagation planning for contests
 - Propagation planning for DXpeditions
 - Cycle 24 Update

The Need for a Model

- After WWII it became apparent that it was important to be able to be on the right frequency at the right time to communicate with a desired location
- A model of the ionosphere was needed
- Ionosondes used to collect data
 - Swept-frequency radar that looks straight up
- Data initially was for 1954-1958
 - Included solar min and solar max
- More data added over the years



Ionomsonde Measurements

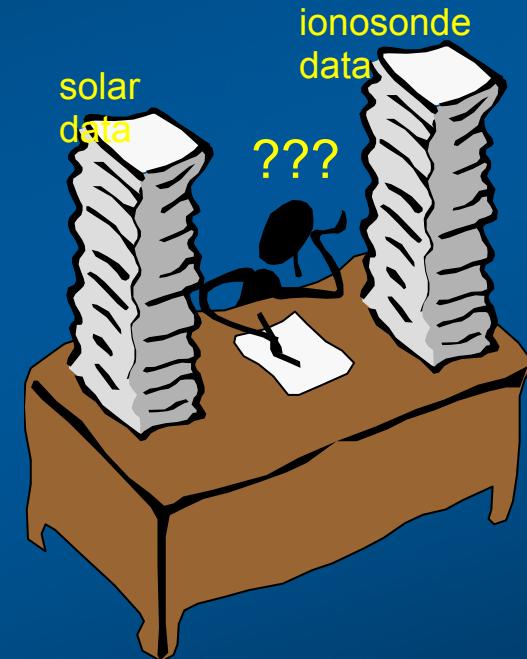


- An ionosonde gives us the critical frequencies and virtual heights of the ionospheric regions
- The data from the ionosonde also gives us the electron density profile (after some math)

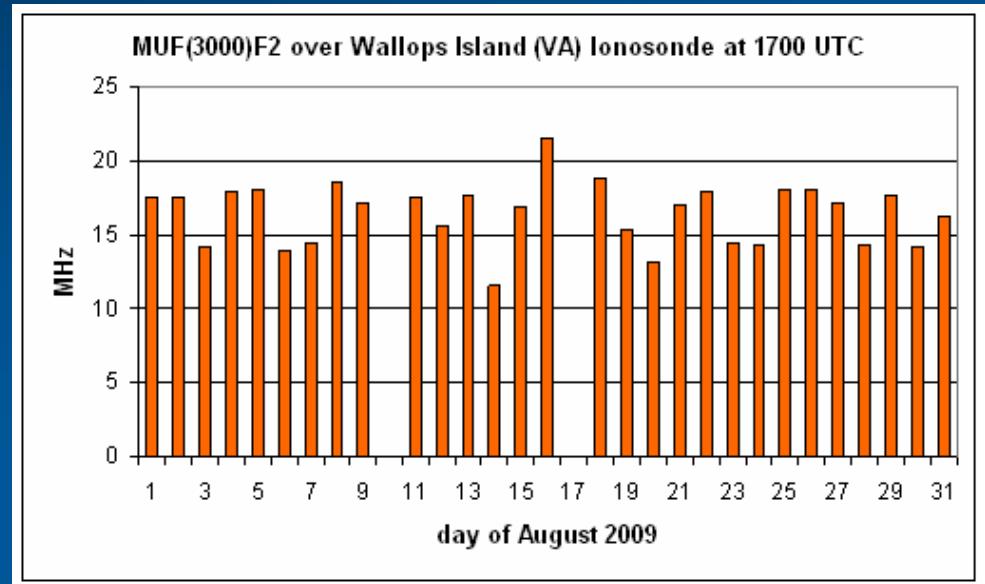
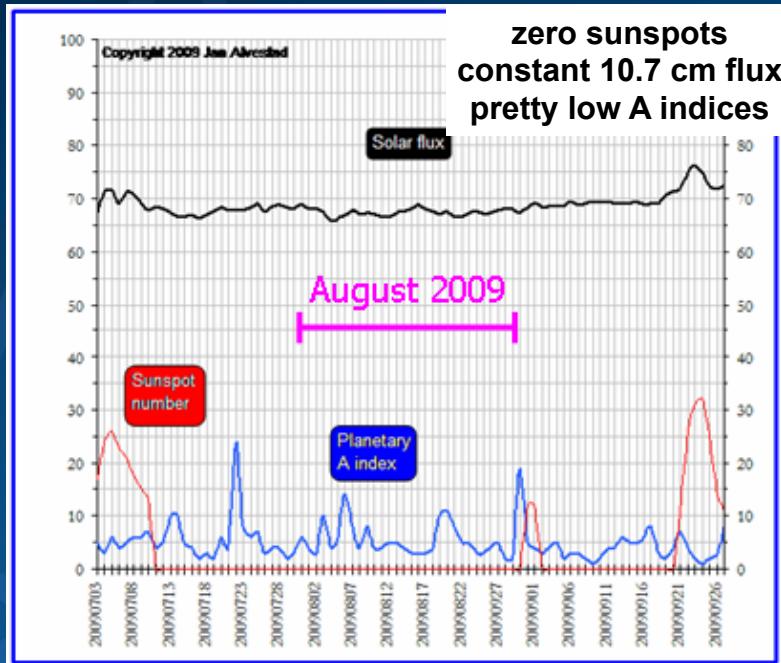


Scientists Began Their Work

- We can predict oblique propagation from critical frequencies and heights using spherical geometry
- Many years of solar data and worldwide ionosonde data collected
- The task of the propagation prediction developers was to determine the correlation between solar data and ionosonde data
- It would have been nice to find a good correlation between what the ionosphere was doing on a given day and what the Sun was doing on the same day



But That Didn't Happen



- Three factors determine ionization
 - Solar radiation (3% of total daytime day-to-day std dev variation)
 - Geomagnetic field activity (13%)
 - Events in lower atmosphere couple up to ionosphere (15%)
- Low of 11.6 MHz on August 14
- High of 21.5 MHz on August 16
- No correlation to daily SF and A

Now What?

Daily correlation not good - the developers were forced to come up with a statistical model over a month's time frame



$R^2 = .0615$

daily correlation

$R^2 = .8637$

monthly median correlation



Monthly median correlation good – smoothed solar flux (or smoothed sunspot number) and monthly median parameters



ARRL
100
YEARS

Interpreting MUF and Sig Str

- Our model of the ionosphere was developed to use a smoothed solar index
 - Smoothed 10.7 cm solar flux or smoothed sunspot number equally good as there is an extremely high correlation between the two
- Our model of the ionosphere was developed to give monthly median MUF and monthly median signal strength
 - Median implies 50% probability
- Using a daily solar index will give results that could be off by a band or two and off by many S-units



Example of Median

- After inputting a smoothed solar index, your favorite software says the MUF is 19.6 MHz and the signal strength is S7 during a specific month at a specific time
 - On half the days of the month, the MUF will be at least 19.6 MHz
 - On any given day during the month the MUF could be up to about 35% lower to about 25% higher
 - On half the days of the month, the sig str will be at least S7
 - On any given day during the month the sig str could be several S-units lower to about an S-unit higher
- Unfortunately, trying to identify which days are 'good' and which days are 'bad' is tough
- For details on downloading VOACAP or W6ELProp and using them and interpreting the results, visit <http://k9la.us>

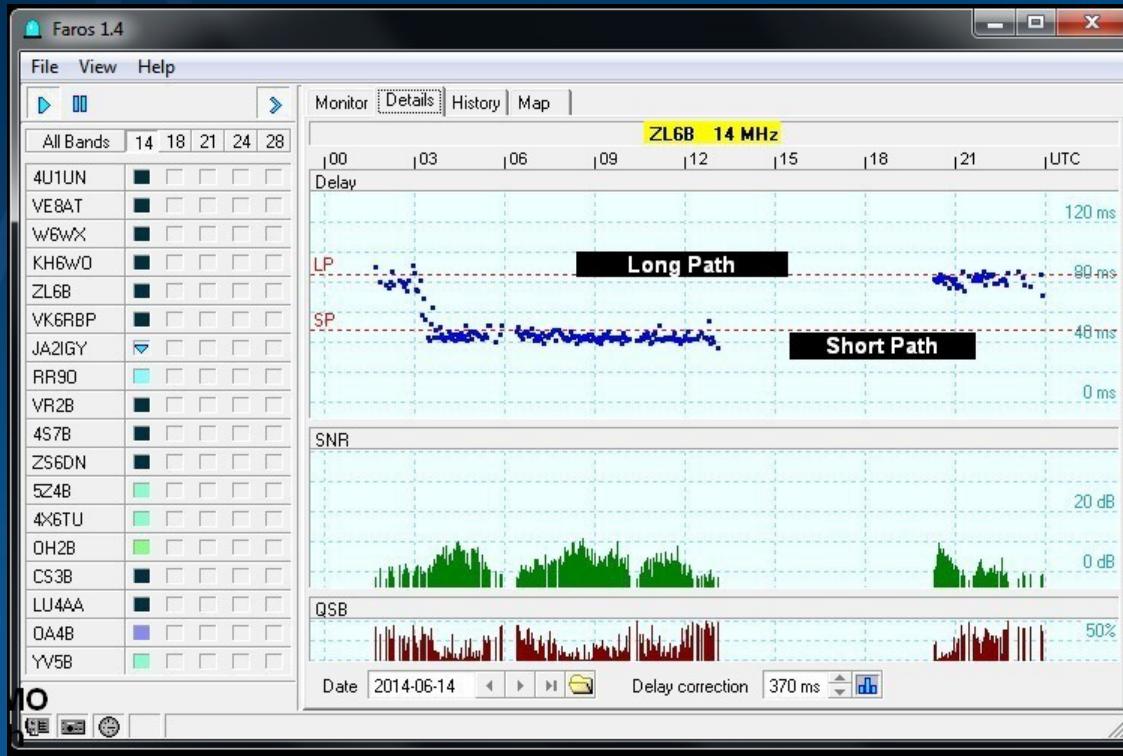


Real-Time Assessment

- For a real-time assessment of propagation, use the IARU/NCDXF beacons on 20m, 17m, 15m, 12m and 10m to give a picture of worldwide propagation
 - <http://www.ncdxf.org/pages/beacons.html>
- Use Faros beacon software to monitor the beacons to study propagation
 - <http://dxatlas.com/faros>
 - Example on the next slide



Faros Results



from K2MO study

- After calibration of the delays, can identify short path and long path
- Measures signal-to-noise ratio
- Can compare to propagation predictions

Might see unusual openings, drop-outs due to geomagnetic field activity and non-great circle paths



Tips about:

- Predictions for individuals
- Propagation planning for contests
- Propagation planning for DXpeditions



K9LA to ZF

| | | | | | | | | | | | | | | |
|------|------|------|------|------|------|------|-----|-----|-----|-----|-----|-----|-----|--------|
| 13.0 | 22.3 | 14.2 | 18.1 | 21.2 | 24.9 | 28.4 | 0.0 | 0.0 | 0.0 | 0.0 | 0.0 | 0.0 | 0.0 | FREQ |
| 1F2 | - | - | - | - | - | - | - | MODE |
| 9.2 | 4.8 | 5.3 | 6.8 | 10.3 | 10.3 | - | - | - | - | - | - | - | - | TANGLE |
| 8.6 | 8.4 | 8.4 | 8.5 | 8.6 | 8.6 | - | - | - | - | - | - | - | - | DELAY |
| 327 | 224 | 235 | 270 | 352 | 352 | - | - | - | - | - | - | - | - | V HITE |
| 0.50 | 0.99 | 0.91 | 0.63 | 0.20 | 0.02 | - | - | - | - | - | - | - | - | MUFday |
| 121 | 113 | 113 | 116 | 137 | 166 | - | - | - | - | - | - | - | - | LOSS |
| 33 | 37 | 37 | 36 | 16 | -12 | - | - | - | - | - | - | - | - | DBU |
| -89 | -82 | -83 | -85 | -107 | -136 | - | - | - | - | - | - | - | - | S DBW |
| -174 | -168 | -171 | -173 | -175 | -177 | - | - | - | - | - | - | - | - | N DBW |
| 85 | 86 | 88 | 88 | 68 | 40 | - | - | - | - | - | - | - | - | SNR |
| -10 | -26 | -26 | -20 | 6 | 34 | - | - | - | - | - | - | - | - | RPWRG |
| 0.96 | 1.00 | 1.00 | 0.99 | 0.84 | 0.35 | - | - | - | - | - | - | - | - | REL |
| 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | - | - | - | - | - | - | - | - | MPROB |
| 0.59 | 0.91 | 0.90 | 0.76 | 0.43 | 0.16 | - | - | - | - | - | - | - | - | S PRB |
| 25.0 | 8.4 | 9.7 | 17.9 | 25.0 | 25.0 | - | - | - | - | - | - | - | - | SIG LW |
| 11.6 | 6.1 | 4.8 | 7.8 | 21.4 | 25.0 | - | - | - | - | - | - | - | - | SIG UP |
| 26.8 | 12.5 | 13.6 | 20.3 | 26.8 | 26.8 | - | - | - | - | - | - | - | - | SNR LW |
| 12.9 | 8.0 | 7.4 | 9.7 | 22.1 | 25.6 | - | - | - | - | - | - | - | - | SNR UP |
| 12.0 | 12.0 | 12.0 | 12.0 | 12.0 | 12.0 | - | - | - | - | - | - | - | - | TGAIN |
| 12.0 | 12.0 | 12.0 | 12.0 | 12.0 | 12.0 | - | - | - | - | - | - | - | - | RGAIN |
| 58 | 74 | 74 | 68 | 42 | 14 | - | - | - | - | - | - | - | - | SNRxx |

VOACAP results

- **October**
- **Fall 2014**
- **1 kW**
- **12 dBi antennas**
- **Req SNR = 13 dB in 3 kHz (90% voice intelligibility)**



What Is the Best Band?

| UTC | median MUF | 20m | 17m | 15m | 12m | 10m | | |
|------|------------|------|------|------|------|------|--------|-----------------------------------|
| 13.0 | 22.3 | 14.2 | 18.1 | 21.2 | 24.9 | 28.4 | | |
| | 0.50 | 0.99 | 0.91 | 0.63 | 0.20 | 0.02 | MUFday | (prob that MUF is at each band) |
| | -89 | -82 | -83 | -85 | -107 | -136 | S DBW | (in dB above 1W – add 30 for dBm) |
| | 85 | 86 | 88 | 88 | 68 | 40 | SNR | (predicted median SNR) |
| | 0.96 | 1.00 | 1.00 | 0.99 | 0.84 | 0.35 | REL | (prob that SNR > requirement) |

- We have two probabilities
 - Probability that MUF is high enough
 - Probability that SNR is high enough
- Multiply them together to get the joint probability (NM7M SK called this DX feasibility) that the MUF and the SNR are simultaneously high enough
- You can also do this with ‘time’ as an additional variable to identify ‘what band at what time’ is best



QRP Considerations

- Fighting a pile-up with QRP can be tough
- Operator skill and antennas play an important role
- Another technique is to identify unusual openings
 - When most others will be in bed
 - Of course the station you're trying to work must be aware of these unusual openings, too!

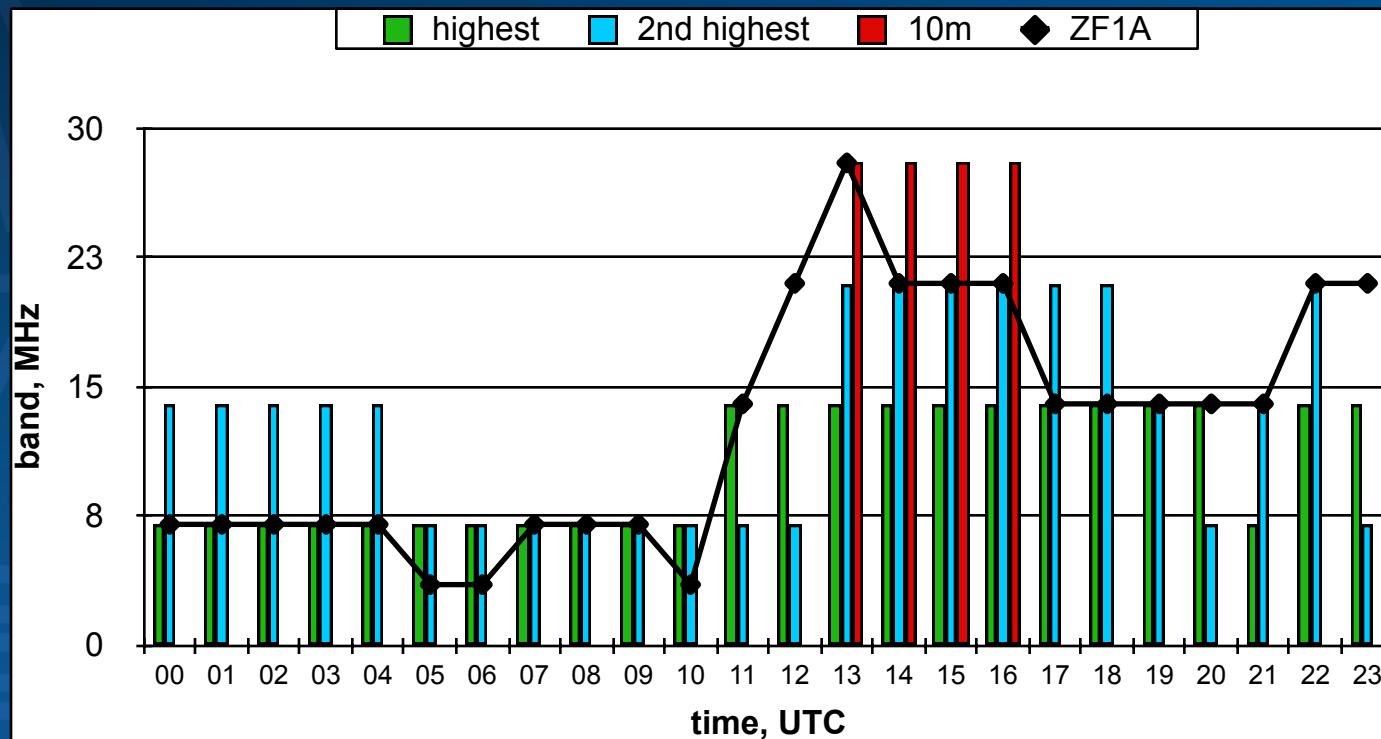


Propagation Planning for Contests

- K1TO, K9MK, W5ASP and I did a Multi-Single contest effort from ZF in CQ WW CW in 1997
 - One 'run' station – work anybody
 - One 'multiplier' station – only work new multipliers
- After the contest I wondered if there was a way to use propagation predictions to tell the best band for the run station to be on for each hour of the contest
- Using the joint probability concept described earlier, I compared our actual ZF band changes to the band changes recommended by VOACAP
 - Only needed to run predictions to NA, EU, JA



VOACAP vs ZF1A



- Decent agreement – use as a guideline
- Method is useful if you are not familiar with propagation from your contest location
 - Experienced contestants likely do not need much help

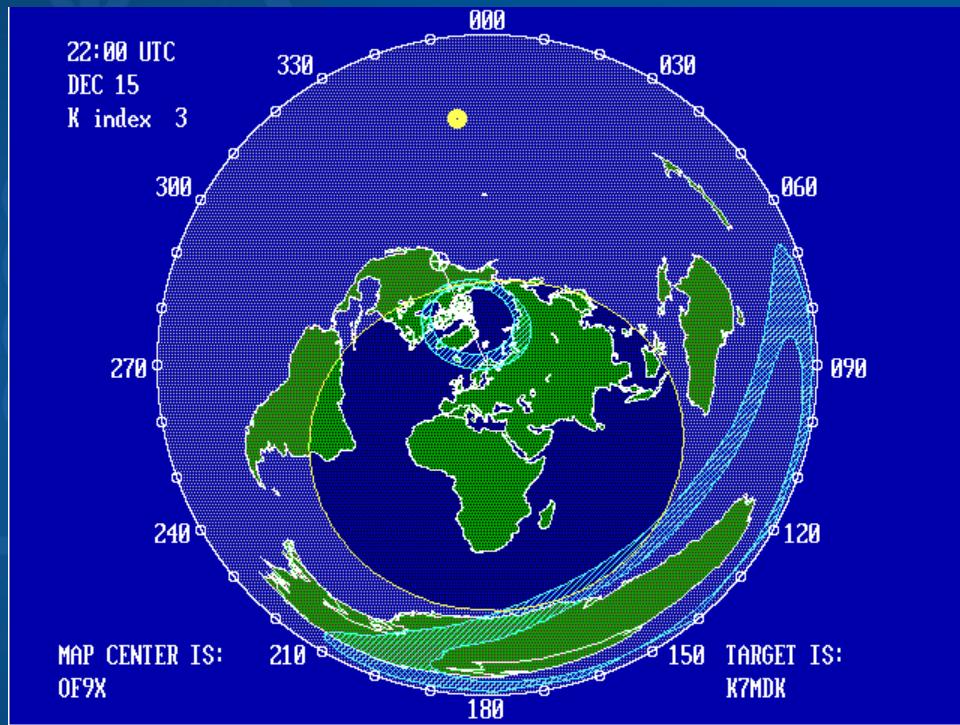
Contesting Tips

- Know the contest
 - Which QSOs are most important to maximize your score
- Run predictions based on the above
 - You don't have to necessarily run predictions to the world
- Be flexible
 - The ionosphere is dynamic
 - Most of the time one band above or below is adequate



Contesting Tips – con't

- Get the big picture
 - Sunspots, 10.7 cm solar flux, A index
 - Great circle map centered on your location
 - Headings, high latitude paths and distances

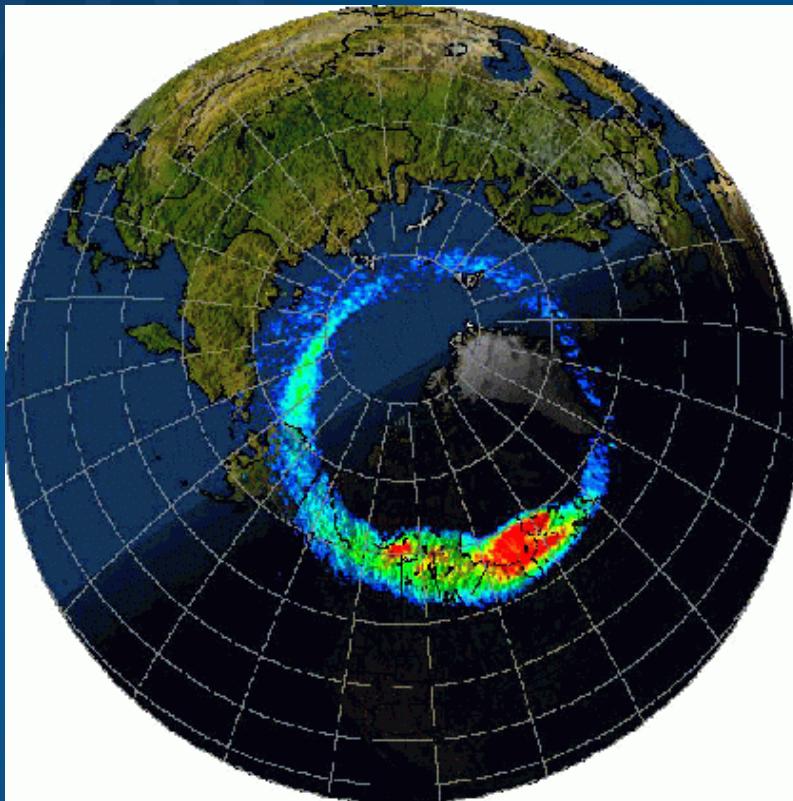


from DXAID V4.5
(old DOS program)



Contesting Tips – con't

- Understand disturbances to propagation



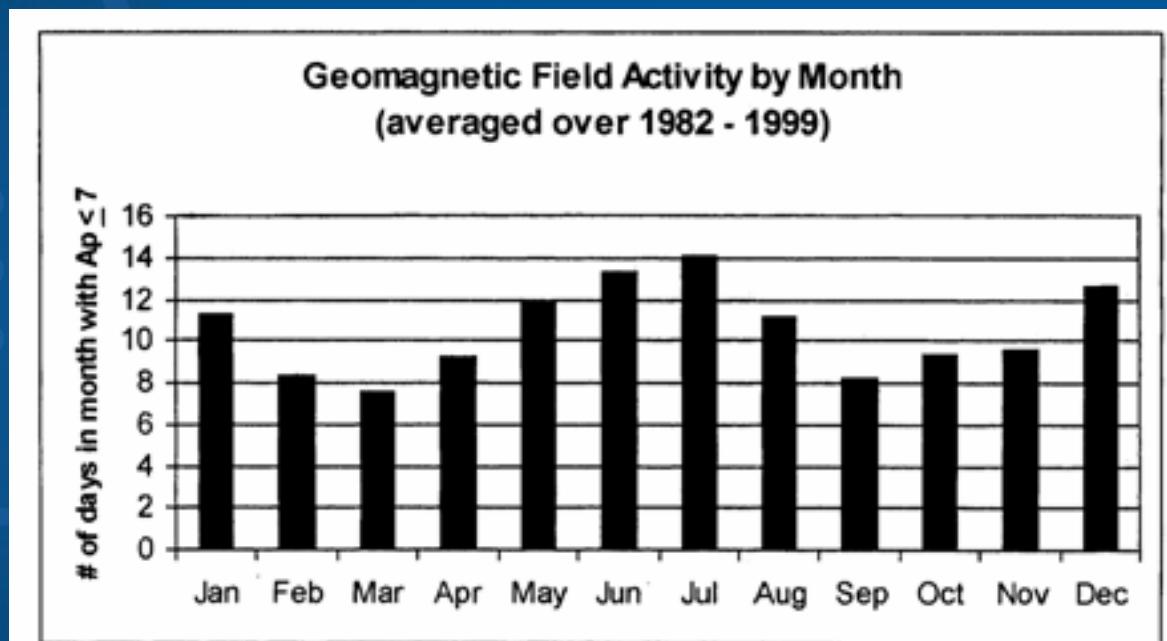
- Geomagnetic storms
 - More than likely, lower worldwide MUFs at mid/high latitudes
 - Possible higher worldwide MUFs at low latitudes
 - Auroral-E propagation
 - VHF propagation at high latitudes
- Solar radiation storms
 - Increased absorption in polar cap
- Radio blackouts
 - Increased absorption on dayside of Earth

Propagation Planning for DXpeditions

- If you are the DXpedition, determine your goal
 - Low bands, high bands or both
- Know where we are in a solar cycle
 - In general, high bands best at max, low bands at min
- Run predictions for potential months
 - Which are best for goals?

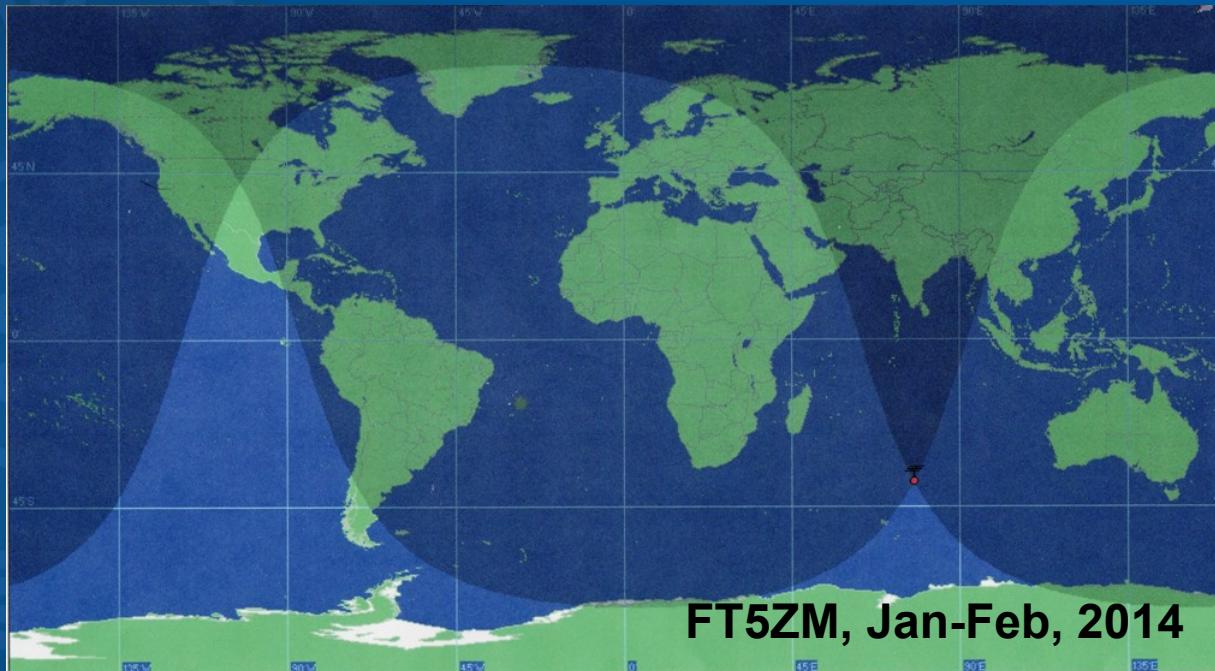
Propagation Planning for DXpeditions – con't

- Know geomagnetic activity by month
 - Least number of days with $Ap \leq 7$ around equinoxes



Propagation Planning for DXpeditions – con't

- Know impact of 'daylight' wedge on 160m
 - Triangular area in which QSOs are likely impossible

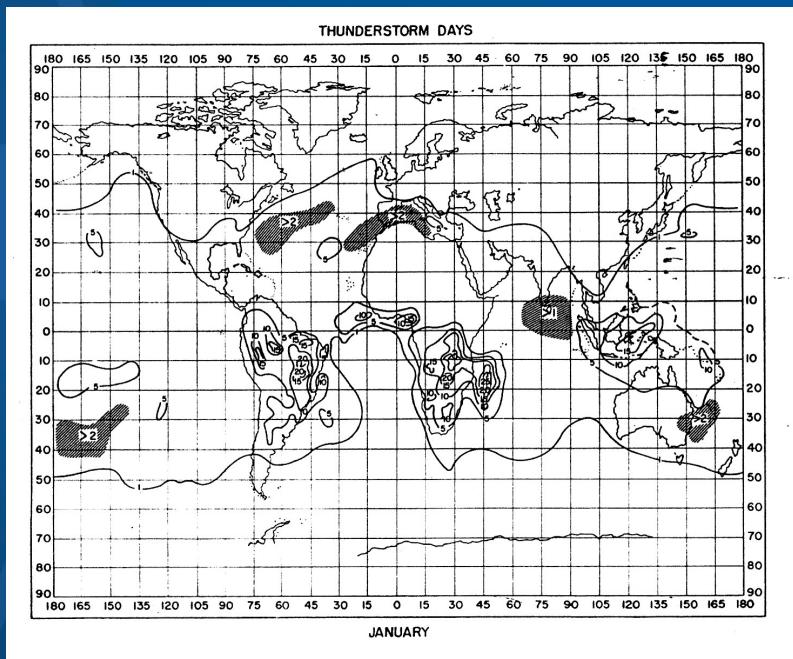


from DX Atlas



Propagation Planning for DXpeditions – con't

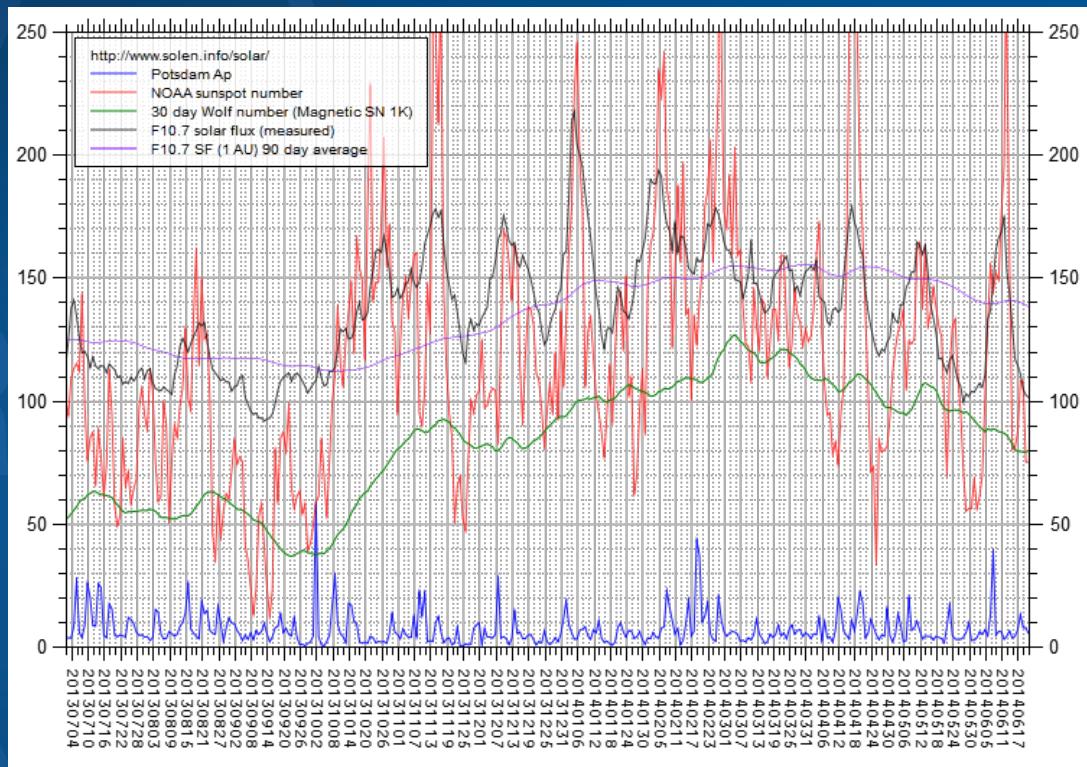
- Have a plan for thunderstorm activity
 - Use maps of thunderstorm activity
 - Use antennas with directivity on lower bands to set null
 - Higher bands not affected as much



from Handbook of
Geophysics, USAF,
1960

DXpedition Tips

- As in propagation planning for contests, get the big picture



<http://www.solen.info/solar/>



DXpedition Tips

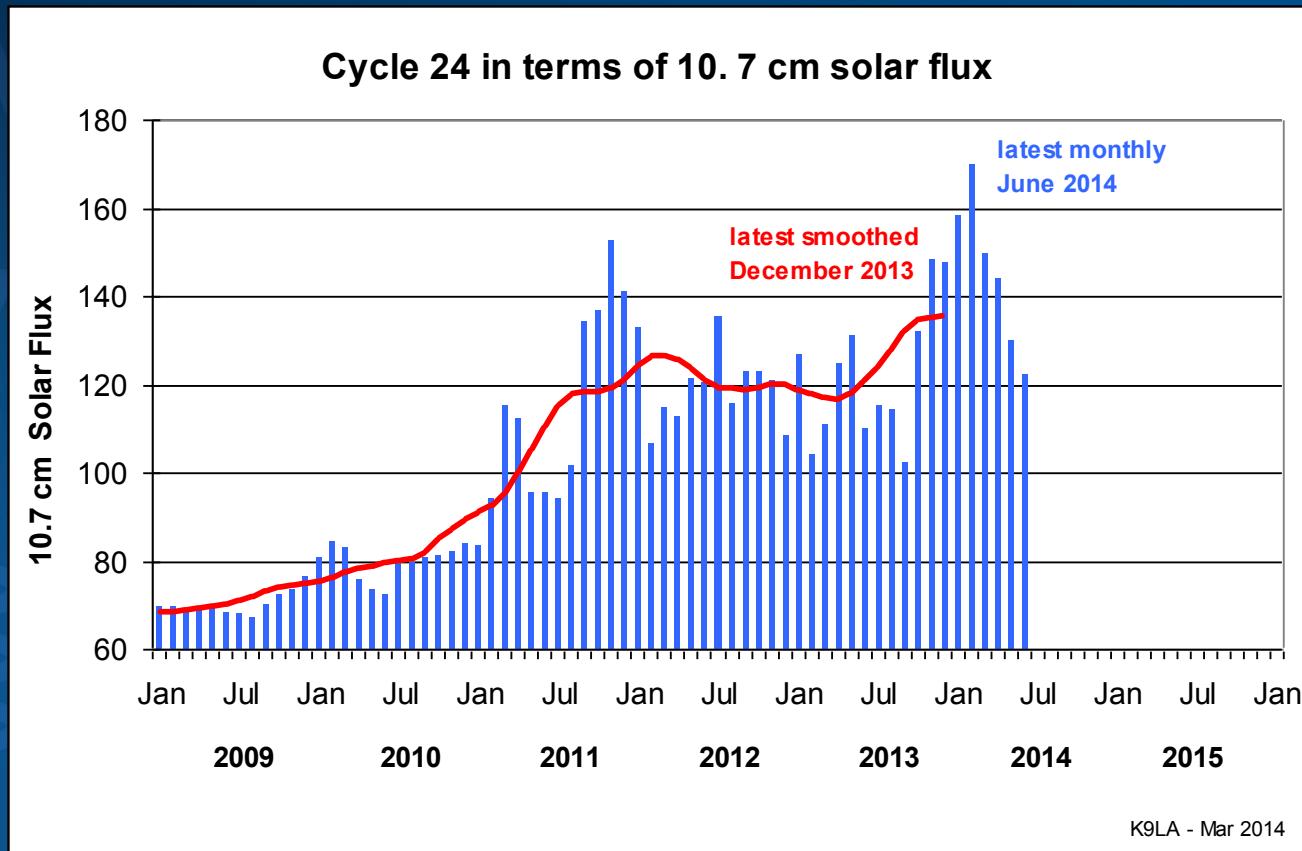
- Try mitigation for disturbances to propagation
 - Move down in frequency for geomagnetic field activity
 - Try long path if PCA affecting polar cap via short path
 - Move up in frequency for radio blackout

Summary

- Ionosphere varies significantly on a day-to-day basis
 - Not well correlated to a daily solar index
 - Explains why you see comments like “the flux **was** high but the higher bands were not good”
- Our predictions are statistical in nature over a month's time frame
 - Use MUF and sig str probabilities to identify best band/time
- In general we want a high solar flux and a low K index
 - 6m F2 requires an extremely high solar flux for an extended period
- Understand disturbances to propagation and mitigation
- In addition to running predictions, look at:
 - The big picture (sunspots, 10.7 cm flux, A index)
 - Great circle paths
 - Thunderstorms
 - 160m ‘daylight wedge’ issue
- Be flexible



Cycle 24 Update



- First peak in early 2012
- Second peak around early 2014
- Higher bands should still be good this fall / winter